

Supplementary Materials

Comprehensive mapping of lunar surface chemistry by adding Chang'e-5 samples with deep learning

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Supplementary Table 1 | The measured chemical contents (wt.%) of six major oxides and their MI locations in 55 lunar sampling sites.

	Lunar sites	TiO ₂	FeO	Al ₂ O ₃	MgO	CaO	SiO ₂	MI location	Sample Numbers	References
Apollo	A12-1	3.35	14.8	14.95	9.55	10.40	45.95	3°0'46.35"S, 23°25'14.33"W	12003, 12070	[1], [2], [3], [4]
	A12-2	3.5	14.3	14.70	9.00	10.70	46.6	3°0'43.74"S, 23°25'22.53"W	12030	[3]
	A12-3	2.9	14.1	15.20	9.40	10.70	46.5	3°1'8.84"S, 23°25'43.81"W	12032	[2], [3], [4]
	A12-4	2.70	14	16.20	8.50	11.00	46.2	3°0'43.02"S, 23°25'36.44"W	12033	[3], [4], [5]
	A12-5	3.50	14.9	15.10	10.20	10.50	44.8	3°1'10.39"S, 23°25'46.30"W	12037	[1], [3], [4]
	A12-6	2.7	14.2	15.40	9.10	10.90	46.8	3°1'13.41"S, 23°25'34.47"W	12041	[3], [4]
	A12-7	3.2	14.1	15.10	9.80	10.90	46.1	3°1'9.67"S, 23°25'20.39"W	12042	[2], [3], [5]
	A12-8	3.6	14.2	14.60	9.50	10.60	46.7	3°1'10.68"S, 23°25'5.76"W	12044	[3]
	A14-LM	1.75	10.53	17.69	9.52	10.32	48.2	3°38'45.40"S, 17°28'18.97"W	14003	[6]
	A14-G	1.68	10.35	17.33	9.58	10.37	48.2	3°38'49.66"S, 17°27'52.39"W	14148,14149,14156	[6]
	A14-Compre	1.7	10.54	17.52	9.41	10.49	48.4	3°38'39.73"S, 17°28'34.13"W	14421	[6]
	A15-LM	1.9	14.38	14.46	9.93	10.47	47.11	26°7'53.29"N, 3°37'58.98"E	15015	[7]
	A15-S1	1.59	15.79	13.15	10.81	10.58	46.73	26°1'51.48"N, 3°36'27.27"E	15071,15081	[8], [9]
	A15-S2	1.32	11.52	17.52	10.54	11.75	46.4	26°0'40.19"N, 3°35'23.21"E	15091,15101,15211,15221,15231	[8], [10], [11], [12]
	A15-S4	1.55	15.96	13.22	11.19	10.46	46.26	26°1'38.80"N, 3°39'41.19"E	15471	[8], [9]
	A15-S6	1.5	12.14	16.47	10.51	11.29	46.72	25°58'43.94"N, 3°40'41.53"E	15241,15261,15271,15291	[11], [12]
	A15-S7	1.25	12.37	15.92	11.17	11.12	46.61	25°59'2.79"N, 3°39'57.19"E	15411,15431	[12], [13], [14]
	A15-S8	1.72	14.98	14.46	10.35	10.38	46.94	26°7'59.05"N, 3°37'51.8"E	15013	[15]
	A15-S9	1.78	16.61	12.54	10.98	10.31	46.46	26°7'47.18"N, 3°34'28.4"E	15501	[9], [16]
	A15-S9a	1.84	20.19	10.31	11.36	9.39	46.52	26°7'45.05"N, 3°33'55.19"E	15531,15601	[8]
	A16-LM	0.6	5.44	26.25	6.25	15.42	45.25	8°58'26.04"S, 15°29'57.86"E	60501,60600,60601	[17], [18], [19]
	A16-S1	0.58	5.4	26.60	6.00	15.60	45	8°58'45.17"S, 15°26'59.87"E	61141,61161,61241,61281,61501	[17], [19], [20]
	A16-S2	0.6	5.5	27.00	6.05	15.70	44.6	8°58'34.77"S, 15°28'19.61"E	62241,62281	[19]
	A16-S4	0.52	4.6	27.60	5.10	16.20	45.1	9°5'39.88"S, 15°30'32.88"E	64421,64501	[18], [19]
	A16-S5	0.65	5.85	26.20	6.25	15.00	45.3	9°4'47.99"S, 15°30'38.76"E	65501,65701,65901	[19], [20]
	A16-S6	0.67	5.95	26.40	6.25	15.70	45.2	9°4'26.00"S, 15°29'51.36"E	66041, 66081	[19]
	A16-S8	0.56	5.35	26.60	6.30	15.30	45.1	9°4'4.27"S, 15°28'35.61"E	68121,68501,68821,68841	[13], [19]
	A16-S9	0.61	5.7	26.30	6.30	15.10	45.2	9°3'10.59"S, 15°28'37.36"E	69921, 69941, 69961	[13], [19]
	A16-S11	0.41	4.2	28.90	4.30	16.50	45.1	8°49'50.31"S, 15°29'35.21"E	67461,67481,67601,67701	[19]
	A16-S13	0.54	4.8	27.60	5.40	15.80	45.1	8°51'15.90"S, 15°30'18.50"E	63321,63341,63501	[13], [19]
	A17-LM	8.5	16.6	12.10	9.80	11.10	40.8	20°11'19.35"N, 30°46'30.89"E	70011,70181	[21], [22], [23], [24]
	A17-S1	9.6	17.8	10.90	9.60	10.80	39.9	20°9'16.24"N, 30°47'23.30"E	71131,71151,71501	[23], [24], [25]
	A17-S2	1.5	8.7	20.70	9.90	12.80	45	20°5'45.86"N, 30°31'49.24"E	72321,72441,72461,72501,72701	[22], [23], [24], [25], [26]
	A17-S3	1.8	8.7	20.40	10.20	12.90	44.9	20°10'10.33"N, 30°33'59.49"E	73211,73221,73261,73281	[21], [22], [24]
	A17-S5	9.9	17.7	10.90	9.60	10.80	39.8	20°11'1.25"N, 30°43'43.39"E	75061,75081	[22], [23], [25], [26]
	A17-S6	3.4	10.7	18.30	10.80	12.20	43.5	20°17'18.36"N, 30°48'14.88"E	76221,76241,76261,76501	[22], [23], [24]
	A17-S7	3.9	11.6	17.30	10.10	11.90	43.7	20°17'21.19"N, 30°49'19.22"E	77511,77531	[23], [24]
	A17-S8	4.3	12.3	16.60	10.20	11.80	43.4	20°16'31.26"N, 30°52'57.17"E	78221,78231,78441,78461	[24], [26]
	A17-S9	6.4	15.4	13.90	10.00	11.30	42.1	20°13'48.79"N, 30°50'12.11"E	79211,79511	[22], [23], [24], [25]
	A17-LRV1	8	16.3	12.60	9.40	11.20	41.3	20°10'21.97"N, 30°41'25.36"E	72131	[24]
	A17-LRV2	4.4	13.4	16.10	10.30	11.90	43.1	20°10'43.41"N, 30°38'54.60"E	72141	[21], [23]
	A17-LRV3	5.5	14.8	14.40	10.40	11.30	42.2	20°10'51.04"N, 30°37'44.49"E	72151,72161	[23], [24]
	A17-LRV4	1.3	8.5	21.40	9.60	12.80	44.9	20°6'26.50"N, 30°32'55.69"E	73151	[24]
	A17-LRV5	2.6	9.8	19.90	8.90	12.80	44.8	20°10'51.77"N, 30°35'35.69"E	74111	[24]
	A17-LRV6	2.6	10.3	19.40	9.90	12.50	44.5	20°11'31.13"N, 30°36'0.89"E	74121	[21], [26]
	A17-LRV7	6.8	16.1	12.80	10.30	10.70	41.8	20°12'46.46"N, 30°40'14.71"E	75111	[24]
	A17-LRV8	6.6	15.7	13.50	9.90	11.30	41.9	20°12'23.07"N, 30°41'42.64"E	75121	[24]
	A17-LRV9	6.1	14.6	14.30	9.80	11.30	42.2	20°13'56.19"N, 30°47'11.17"E	76121	[24]
	A17-LRV10	3.7	11.2	17.50	10.50	12.10	43.5	20°16'48.31"N, 30°47'37.59"E	76131	[24]
	A17-LRV11	4.5	12.7	16.30	10.00	11.90	43.2	20°16'21.57"N, 30°52'32.96"E	78121	[24]
	A17-LRV12	10	17.4	11.20	9.36	10.80	39.9	20°11'41.55"N, 30°48'58.56"E	70311,70321	[24]
Luna	L16	3.41	16.85	15.24	8.67	12.38	41.5	0°30'53.84"S, 56°21'52.57"E	luna16s	[27], [28]
	L20	0.46	7.5	22.90	9.15	14.50	45.6	3°47'10.84"N, 56°37'27.00"E	luna20s	[19]
	L24	1.1	20.5	11.10	10.20	11.00	45.4	12°42'51.11"N, 62°12'49.79"E	24077,24109,24149,24174,24182,24210	[29]
Chang'e	CE-5	5	22.5	10.80	6.48	11.00	42.2	43°3'28.80"N, 51°54'57.60"W	CE-5	[30]

The chemical contents presented here are the average values of the measured lunar surface chemical abundances from the returned samples. References: [1]Wanke *et al.*,(1971); [2]Morrison *et al.*,(1971); [3]Frondel *et al.*,(1971); [4]Korotev and Rockow(1995); [5]Cuttlm *et al.*,(1971); [6]Philpotts *et al.*,(1972); [7]Scoon *et al.*,(1977); [8]Wanke *et al.*,(1973); [9]Duncan *et al.*,(1975); [10]Carron *et al.*,(1972); [11]Cutlitta *et al.*,(1973); [12]Korotev(1987a); [13]Willis *et al.*,(1972); [14]Rose *et al.*,(1975); [15]Christian *et al.*,(1976); [16]Apollo 15 Preliminary Examination Team(1972); [17]Rose *et al.*,(1973); [18]Apollo 16 Preliminary Examination Team(1973); [19]Korotev(1981); [20]Korotev(1982); [21]Wanke *et al.*,(1974); [22]Rose *et al.*,(1974); [23] Rhodes *et al.*,(1974); [24]Korotev and Kremser(1992); [25]Laul *et al.*,(1974); [26]Duncan *et al.*,(1974); [27]Gillum *et al.*,(1972); [28]Vingradov(1971); [29] Blanchard *et al.*,(1978); [30]Li *et al.*,(2021).

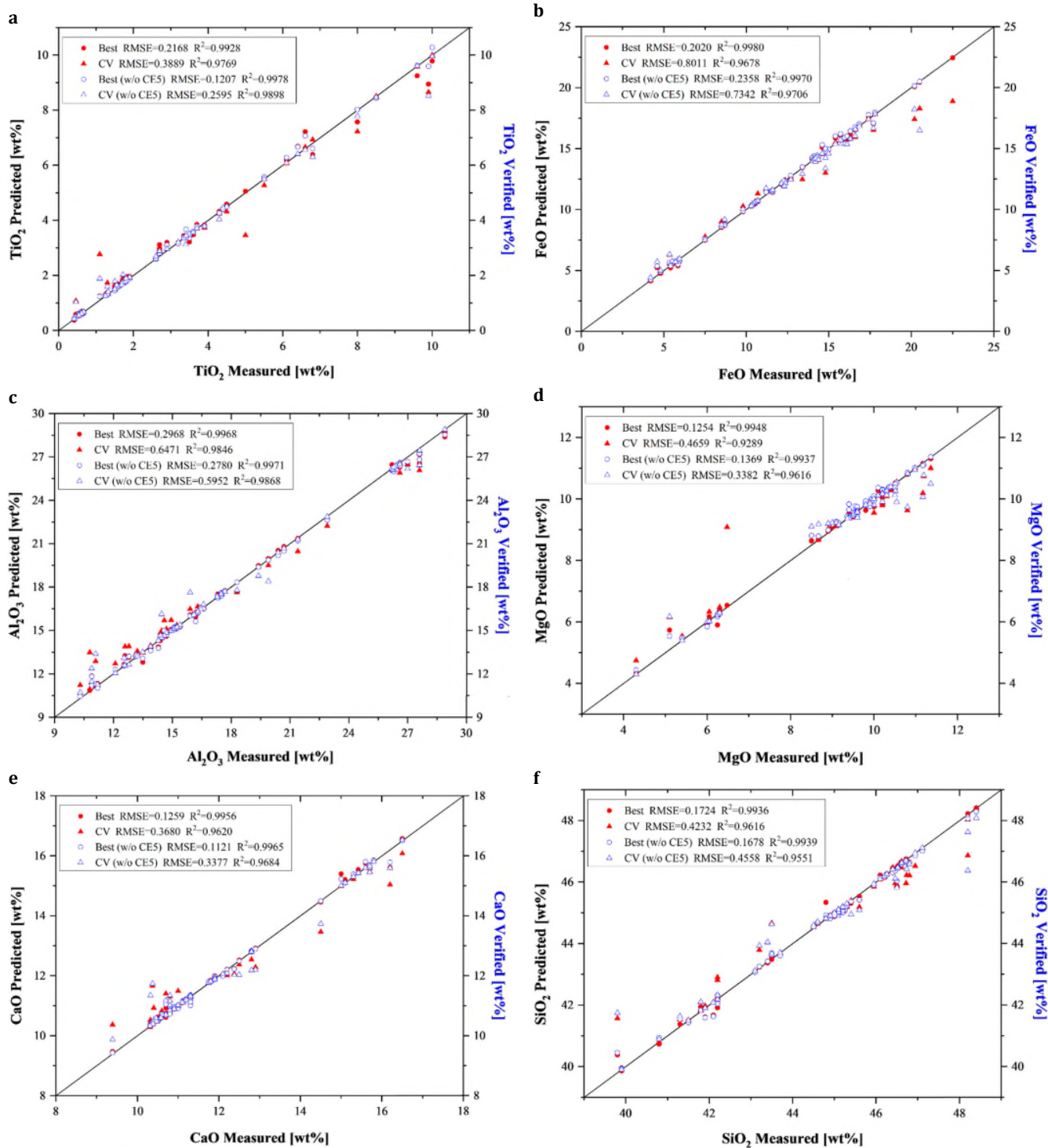
Supplementary Table 2 | The accuracies (RMSEs) of inversion results from the existing sampling points.

Data	Methods	TiO ₂	FeO	Al ₂ O ₃	MgO	CaO	SiO ₂
Clementine UVVIS	Lucey et al. ⁵	1.3076	0.9047	-	-	-	-
LP GRS	Prettyman et al. ³⁵	1.8915	4.0542	5.4899	-	-	-
Chang'e-1 IIM	Xia et al. ⁹	0.3156	0.1993	0.4020	0.1521	0.3839	0.5495
Diviner CF	Ma et al. ²²	2.0293	2.4107	2.1391	0.7206	0.9925	1.9851
	Wang et al. ²⁴	0.9861	1.1486	0.9866	0.4337	0.7299	-
SELENE MI	Zhang et al. ²³	0.2617	0.3576	0.4447	0.2193	0.3814	-
	This work	0.2168	0.2020	0.2968	0.1254	0.1259	0.1724

For abundance algorithms by Lucey et al., only TiO₂ and FeO abundances were derived. In inversion methods with SELENE MI by Wang et al. and Zhang et al., five oxides abundances were estimated. The sampling points applied in different literatures of Clementine UVVIS, Diviner CF, Chang'e-1 IIM, SELENE MI (Wang et al. and Zhang et al.) are used for calculating RMSEs. This work and LP GRS use the same sample points. All the measured and predicted values can be found in Supplementary Fig. 4.

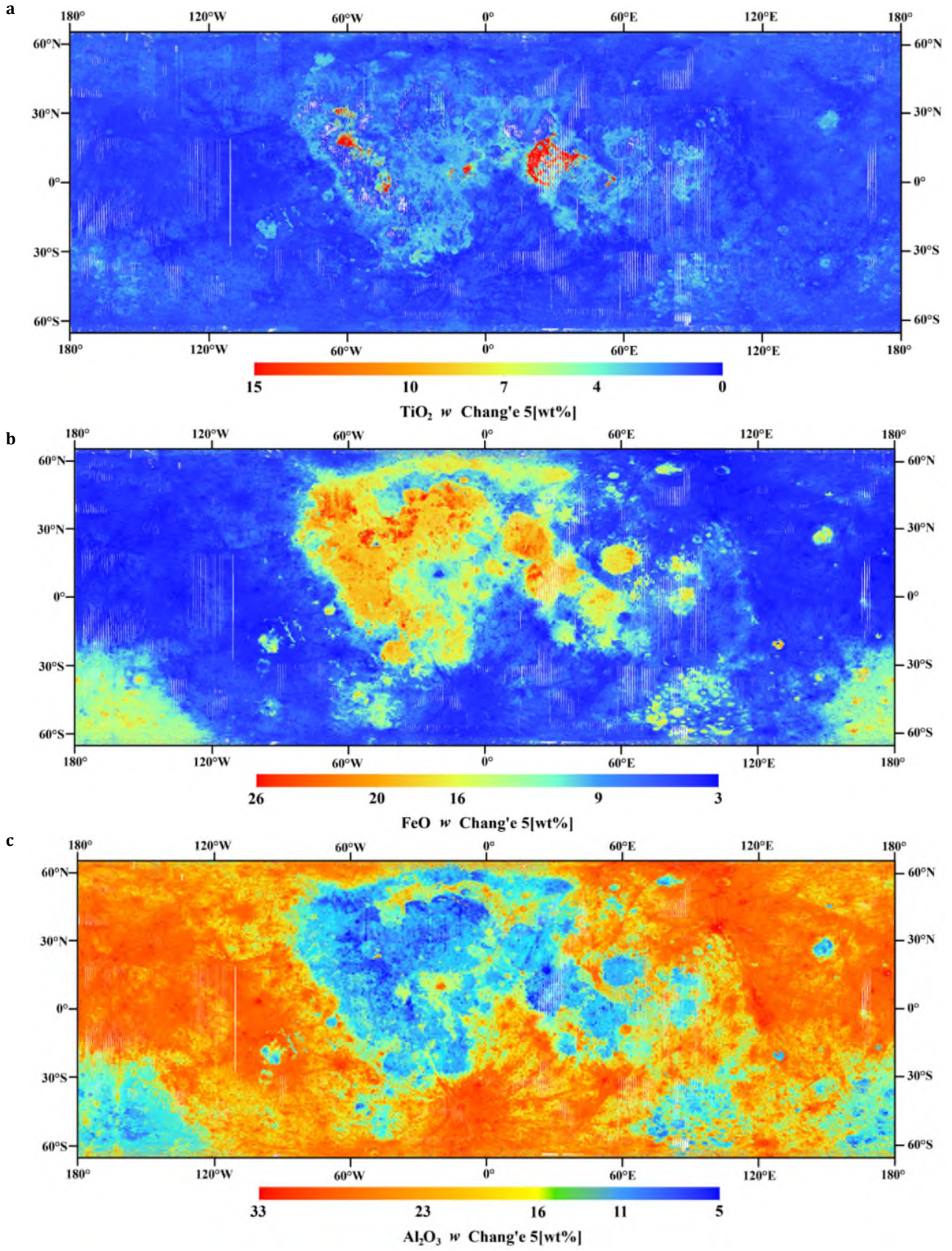
Supplementary Table 3 | The performance (RMSEs and R^2) of the designed 1D CNN inversion model and the compared ELM and GOP methods under the LOOCV configuration.

Methods	TiO ₂		FeO		Al ₂ O ₃		MgO		CaO		SiO ₂	
	RMSE	R^2	RMSE	R^2	RMSE	R^2	RMSE	R^2	RMSE	R^2	RMSE	R^2
ELM ⁶⁴	1.0411	0.8348	1.0103	0.9488	1.3146	0.9366	0.86	0.7576	0.6102	0.8956	1.0303	0.7721
GOP ⁶⁵	1.6472	0.5864	2.2764	0.7400	1.8158	0.8790	0.8914	0.7395	0.8938	0.7761	0.9937	0.7880
This work	0.3889	0.9769	0.8011	0.9678	0.6471	0.9846	0.4659	0.9289	0.3680	0.9620	0.4232	0.9616



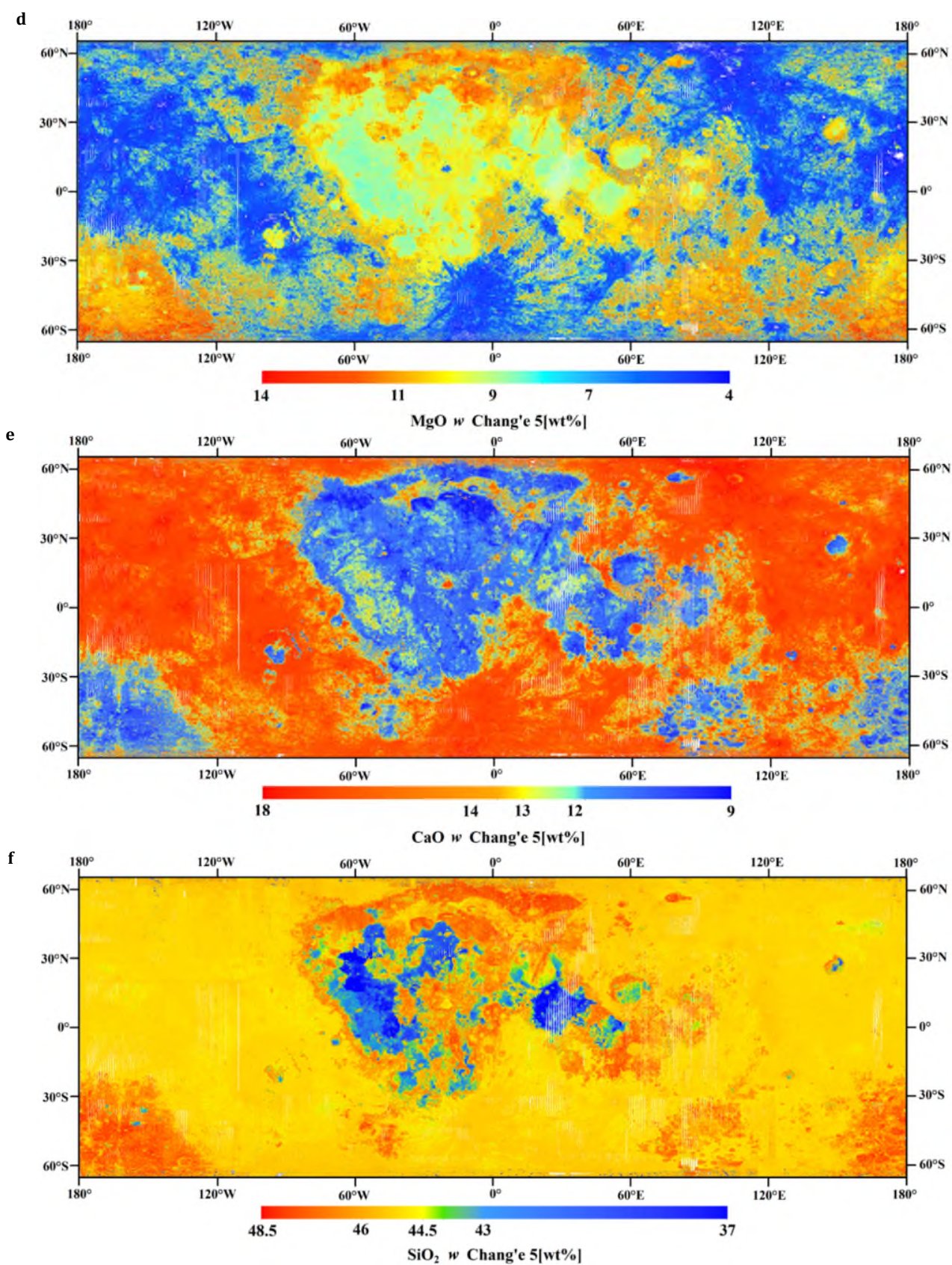
Supplementary Fig. 1 | Prediction and validation accuracies of the deep learning-based inversion model. a, b, c, d, e, f and g show the accuracy evaluation of the six major oxides TiO_2 , FeO , Al_2O_3 , MgO , CaO and SiO_2 , respectively. For the 55 sampling points from Apollo, Luna and Chang'e-5 mission, the red dots and triangles indicate the best prediction precisions by

the 1D convolutional neural network-based inversion algorithm and the validation accuracies by leave-one-out-cross-validation. The blue dots and triangles are the best prediction and validation results without Chang'e-5 sample.



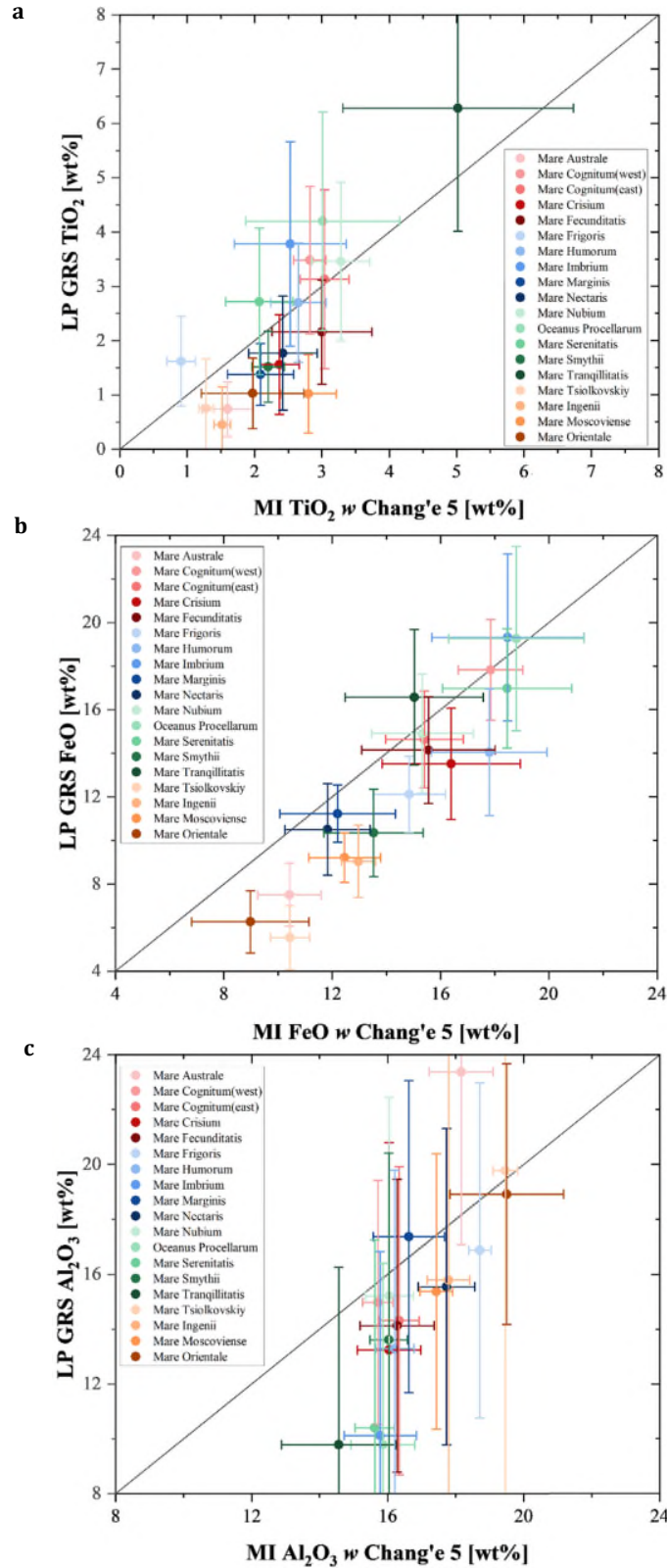
Supplementary Fig. 2 | The lunar surface chemistry maps of six major oxides only with Apollo and Luna data (w/o CE5). a, b and c show the

maps of TiO_2 , FeO and Al_2O_3 abundances calculated from the deep learning-based inversion method, respectively.



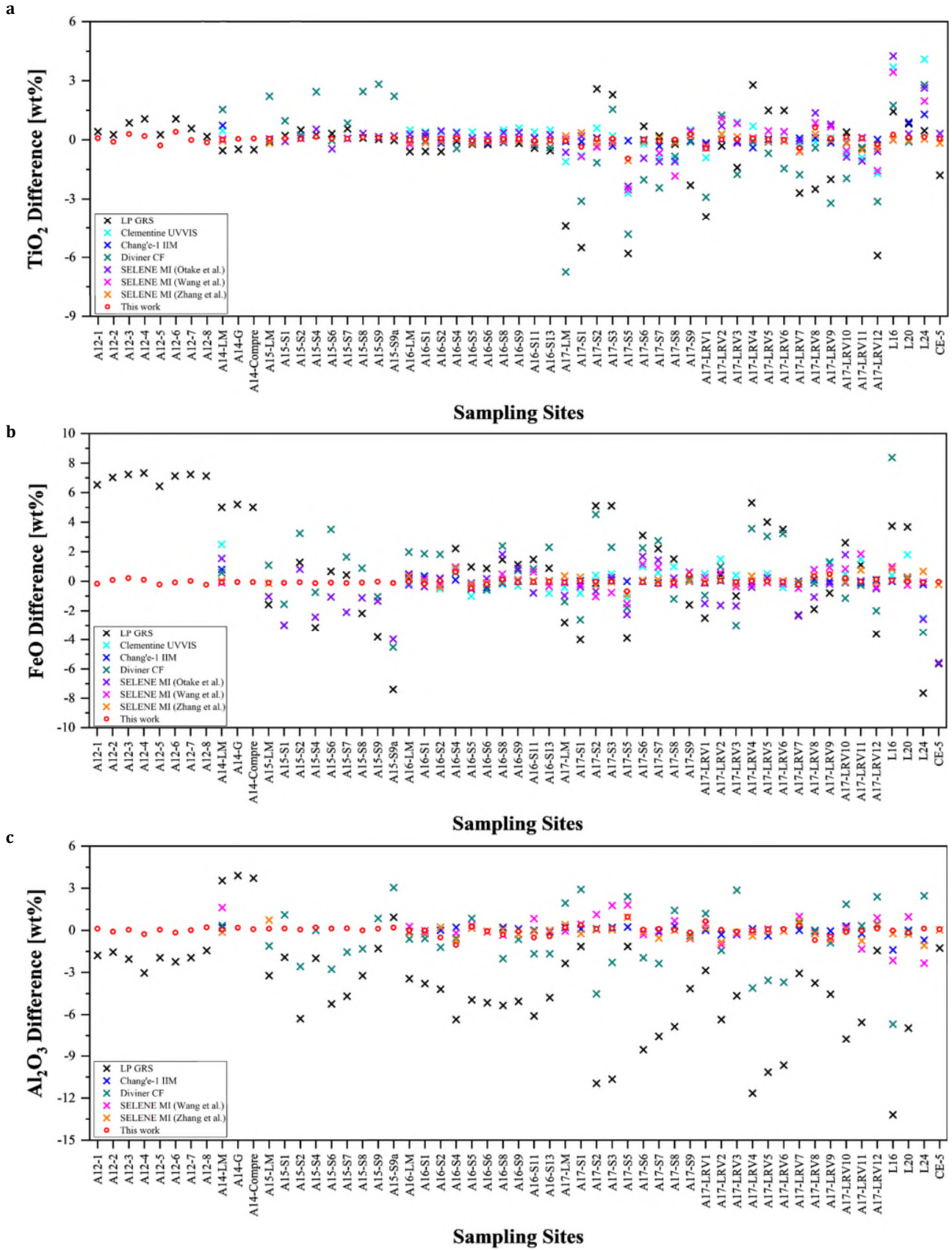
Supplementary Fig. 2 | The lunar surface chemistry maps of six major oxides only with Apollo and Luna data (w/o CE5) (*continued*). d, e and f

show the maps of MgO, CaO and SiO₂ abundances calculated from the deep learning-based inversion method, respectively.

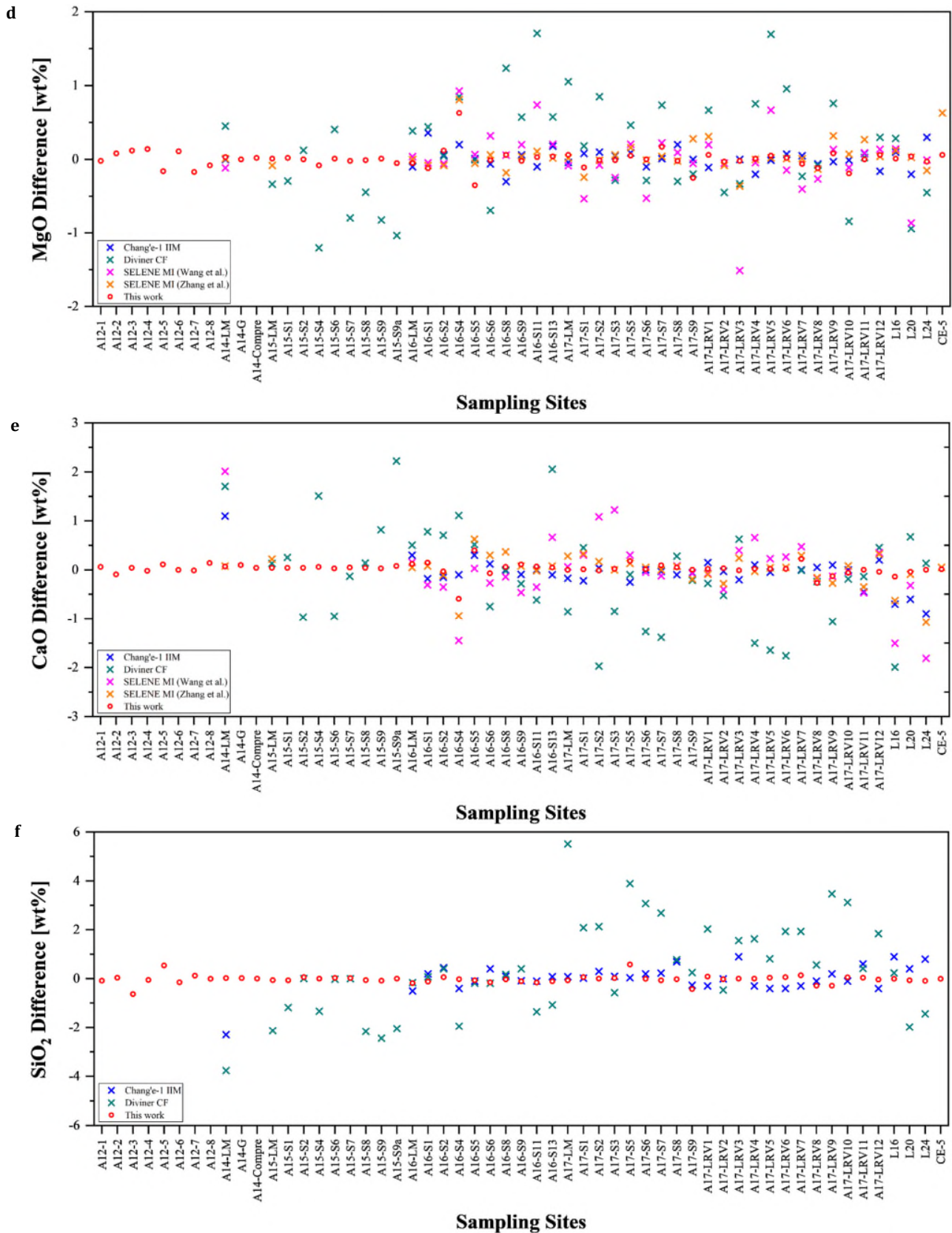


Supplementary Fig. 3 | The contents of the six major oxides from Lunar Prospector (LP) Gamma-Ray Spectrometer (GRS) maps and SELENE (KAGUYA) multiband imager (MI) data by adding Chang'e 5 samples for lunar mare. a, b and c show the averaged TiO_2 , FeO and Al_2O_3 contents,

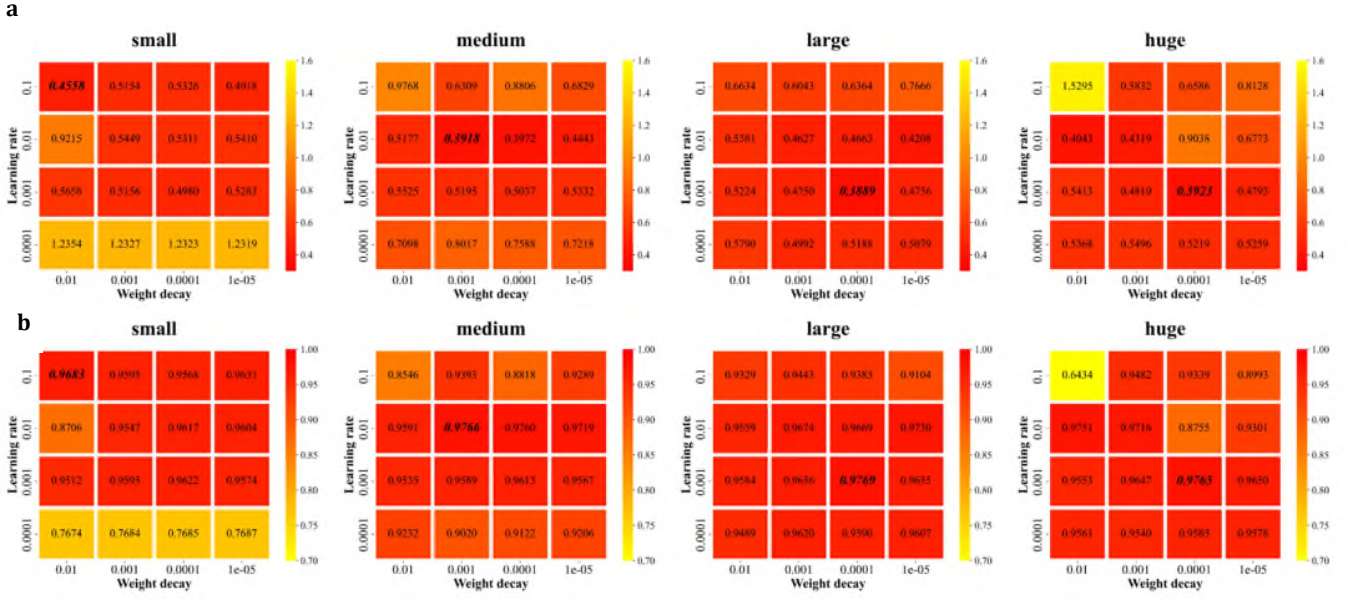
respectively. Error bar indicates the standard deviation. Solid line is the 1:1 line. The Chang'e-5 inversion maps were down-sampled to the LP GR resolution, i.e., 2° per pixel.



abundance algorithms or inversion methods, respectively. The Apollo 11 results were not extracted from the Chang'e-1 IIM images and Apollo 15 values were not involved in the SELENE MI for data missing.



results were not extracted from the Chang'e-1 IIM images and Apollo 15 values were not involved in the SELENE MI for data missing.



Supplementary Fig. 5 | Ablation experimental results of the 1D CNN inversion model with different model size, learning rate and weight decay on TiO₂. **a** shows the RMSE results under the LOOCV setup, where the four subplots represent the ablation experiment results at all combinations of learning rates and weight decay for the 'small', 'medium', 'large' and

'huge' models, in which smaller RMSE values indicate better performance. The best RMSE is bolded and italicized. **b** shows the R² results with the same settings as **a**, except that smaller R² values indicate better performance. The best R² is bolded and italicized.